

Flight and Orbital Mechanics

Solutions

Solutions exam Flight and Orbital Mechanics (November 2011)

Question 1

F

Question 2

D

Question 3

a and b.

Equations of motion
Horizontal sustained turn

Free Body Diagram

Kinetic Diagram

Eq. of motion

$$\vec{F} = m\vec{a}$$

$$0 = T - D$$

$$\frac{WV^2}{gR} = L \sin \mu$$

$$0 = L \cos \mu - W$$

$T = D$

$\frac{WV^2}{gR} = L \sin \mu$

$L \cos \mu = W$

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c.

$$T = D$$

$$L = nW$$

$$D = D \frac{L}{L} = \frac{D}{L} nW = \frac{C_D}{C_L} nW$$

$$T = \frac{C_D}{C_L} nW$$

$$n_{\max} = \frac{T_{\max}}{W} \frac{C_L}{C_D}$$

d.

$$V = M \cdot a = M \sqrt{\gamma RT} = 0.25 \sqrt{1.4 \cdot 287.05 \cdot 288.15} = 85.1 \text{ [m/s]}$$

$$L = n_{\max} W$$

$$V = \sqrt{\frac{n_{\max} W \frac{2}{S} \frac{1}{\rho C_L}}{\frac{T_{\max} C_L W}{W C_D \frac{2}{S} \frac{1}{\rho C_L}}}} = \sqrt{\frac{T_{\max} \frac{2}{S} \frac{1}{\rho C_D}}{\frac{T_{\max} \frac{2}{S} \frac{1}{\rho \left(C_{D_0} + \frac{C_L^2}{\pi A e} \right)}}}}$$

$$P_{\max} = T_{\max} V$$

$$1000000 = T_{\max} \cdot 85.1 \Rightarrow T_{\max} = 11755 \text{ [N]}$$

$$A = \frac{b^2}{S} = \frac{14.6^2}{23.3} = 9.15$$

$$V = \sqrt{\frac{T_{\max} \frac{2}{S} \frac{1}{\rho \left(C_{D_0} + \frac{C_L^2}{\pi A e} \right)}}}$$

$$85.1 = \sqrt{\frac{11755 \frac{2}{23.3} \frac{1}{1.225 \left(0.04 + \frac{C_L^2}{\pi \cdot 9.14 \cdot 0.6} \right)}}} \Rightarrow C_L = 1.13$$

e.

$$\left. \begin{array}{l} L \cos \mu = W \\ L \sin \mu = \frac{W V^2}{g R} \\ \sin^2 \mu + \cos^2 \mu = 1 \end{array} \right\} \Rightarrow R = \frac{V^2}{g \sqrt{n^2 - 1}}$$

$$C_L = 1.13$$

$$C_D = 0.04 + 0.058 \cdot 1.13^2 = 0.114$$

$$n = \frac{11755 \cdot 1.13}{100000 \cdot 0.114} = 1.17$$

$$R = \frac{V^2}{g \sqrt{n^2 - 1}} = \frac{85.1^2}{9.80665 \sqrt{1.17^2 - 1}} = 1235 \text{ [m]}$$

$$T_{\pi} = \frac{\pi R}{V} = \frac{\pi \cdot 1235}{85.1} = 46 \text{ [s]}$$

Question 4

a.

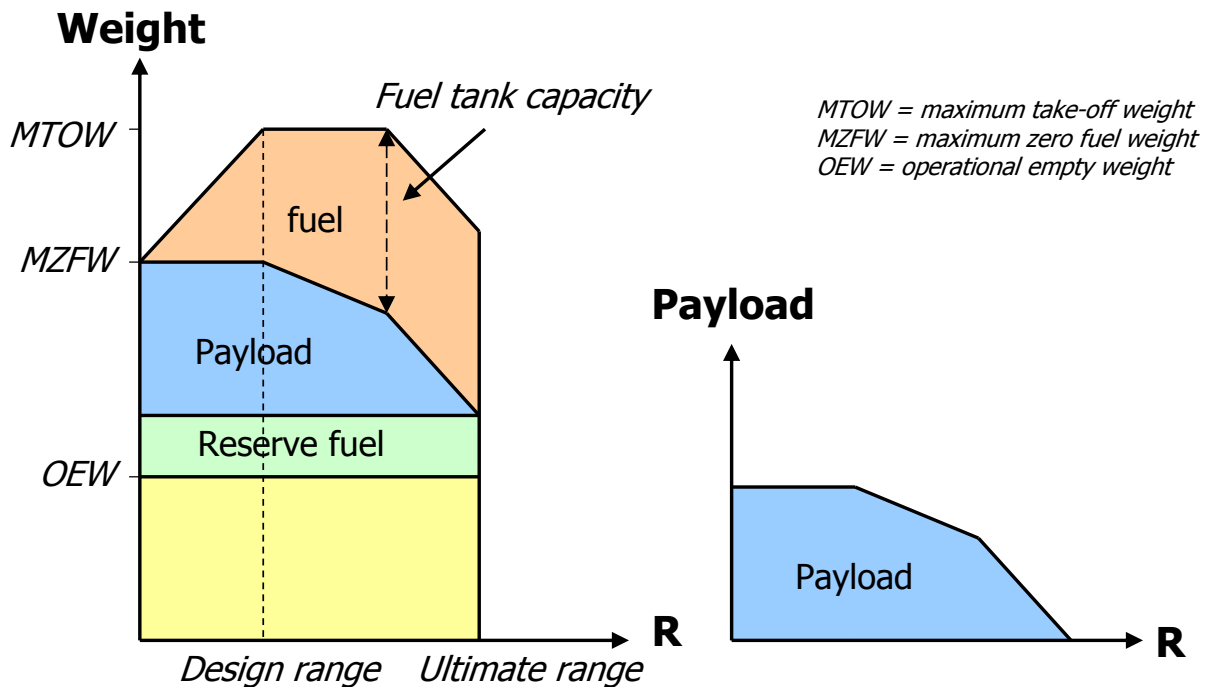
$$R = \int_{t_0}^{t_1} V dt \quad \left. \begin{array}{l} \\ \\ \frac{dW}{dt} = -F \Rightarrow dt = -\frac{dW}{F} \end{array} \right\} R = \int_{W_1}^{W_0} \frac{V}{F} dW$$

$$R = \int_{W_1}^{W_0} \frac{V}{F} dW = \int_{W_1}^{W_0} \frac{V}{c_T T} dW = \int_{W_1}^{W_0} \frac{V C_L}{c_T C_D} \frac{1}{W} dW$$

Angle of attack, c_T and airspeed are constant, hence:

$$R = \frac{V C_L}{c_T C_D} \int_{W_1}^{W_0} \frac{1}{W} dW = \frac{V C_L}{c_T C_D} \ln \left(\frac{W_0}{W_1} \right)$$

B



Question 5:

The ESA satellite GOCE (altitude 250 km, circular orbit) observes the gravity field of the Earth at this very moment. One of the elements of interest is the term $J_{3,1}$. The gravity potential of the Earth is given by the following equation:

$$U = -\frac{\mu}{r} \left[1 - \sum_{n=2}^{\infty} J_n \left(\frac{R_e}{r} \right)^n P_n(\sin \delta) + \sum_{n=2}^{\infty} \sum_{m=1}^n J_{n,m} \left(\frac{R_e}{r} \right)^n P_{n,m}(\sin \delta) \cos(m(\lambda - \lambda_{n,m})) \right]$$

Here, $P_n(\sin\delta)$ and $P_{n,m}(\sin\delta)$ represent the Legendre polynomials and functions, respectively:

$$P_n(x) = \frac{1}{(-2)^n n!} \frac{d^n}{dx^n} (1-x^2)^n$$

$$P_{n,m}(x) = (1-x^2)^{m/2} \frac{d^m P_n(x)}{dx^m}$$

- (2 points) What do the parameters J_n , $J_{n,m}$ and $\lambda_{n,m}$ represent?
- (2 points) How are the parameters r , δ and λ defined?
- (2 points) Give the general expression to derive the radial acceleration from the potential formulation for the gravity field.
- (6 points) Derive the general equation for the radial acceleration due to the term $J_{3,1}$ for an arbitrary satellite (i.e., without substituting any numbers).
- (3 points) What is the equation for the radial acceleration due to $J_{3,1}$ as experienced at GOCE' altitude (expressed in numbers, still for arbitrary latitude and longitude)?
- (3 points) What are the values for δ for which this radial acceleration changes sign?

Data: $\mu_{\text{Earth}} = 398600.4415 \text{ km}^3/\text{s}^2$; $R_e = 6378.137 \text{ km}$; $J_{3,1} = -1.72 \times 10^{-6}$; $\lambda_{3,1} = -1.0^\circ$.

Answers:

- J-terms are scaling factors of the irregularities in the gravity field; $\lambda_{n,m}$ defines the orientation of a particular irregularity
- r = radial distance between position of interest and center of Earth; δ = latitude w.r.t. Earth equator; λ = longitude w.r.t. Earth-fixed reference meridian
- $a_r = -\partial U/\partial r$
- $a_r = -4 \mu J_{3,1} R_e^3 r^{-5} \cos\delta ((15/2)\sin^2\delta - (3/2)) \cos(\lambda - \lambda_{3,1})$
- $a_r = 5.562 \cdot 10^{-5} \cos\delta ((15/2)\sin^2\delta - (3/2)) \cos(\lambda + 1.0) \text{ [m/s}^2\text{]}$
- zero values for $\delta = -90^\circ$ and $+90^\circ$ -> do not play a role. Other zero values for $\delta = -26.6^\circ$ and $\delta = +26.6^\circ$

Question 6:

Consider a sample return mission to an asteroid (i.e., from Earth to the asteroid and then back to Earth again).

- (3 points) Compute the 1-way travel time for a Hohmann transfer from Earth to the asteroid.
- (2 points) The synodic period can be computed with the following equation:

$$\frac{1}{T_{syn}} = \left| \frac{1}{T_1} - \frac{1}{T_2} \right|$$

What is the definition of a synodic period?

- (2 points) Compute the synodic period for the system Sun-Earth-asteroid.
- (2 points) Compute the mean motion of Earth around the Sun (normally labeled "n", here labeled ω_E), and of the asteroid around the Sun (labeled ω_{ast}) (both in [rad/s]).

- e) (2 points) The following general equations give the total round-trip time T and the time to be spent at the asteroid t_{stay} (provided we do both transfers by means of a Hohmann trajectory, and the departure and target objects orbit the Sun in circular orbits):

$$t_{stay} = T - 2T_H = \frac{2\pi(N+1) - 2\omega_E T_H}{\omega_E - \omega_{ast}}$$

$$T = \frac{2\pi(N+1) - 2\omega_{ast} T_H}{\omega_E - \omega_{ast}}$$

Discuss the meaning of the two equations and the parameter N (i.e., give a physical interpretation).

- f) (3 points) What are the values for the parameter N , the stay time, and the total trip time for a mission to this asteroid?
- g) (2 points) Can we reduce the total trip time? If so, how? A qualitative answer is sufficient.
- Data: $\mu_{Sun} = 1.3271 \times 10^{11} \text{ km}^3/\text{s}^2$; distance Earth-Sun = 1 AU; distance asteroid-Sun = 0.95 AU; 1 AU = $149.6 \times 10^6 \text{ km}$.

Answers:

- a) $a_{tr} = (a_E + a_{ast})/2 = 0.975 \text{ AU}$; $T = \pi\sqrt{a_{tr}^3/\mu} = 15191560 \text{ s} = 175.83 \text{ days}$
- b) time interval after which relative geometry of 2 objects (planets, asteroids) repeats
- c) $T_E = 2\pi\sqrt{a_E^3/\mu} = 31559160 \text{ s} = 365.27 \text{ days}$; $T_{ast} = 29222061 \text{ s} = 338.22 \text{ days}$; $T_{syn} = 3.946 \cdot 10^8 \text{ s} = 4567.15 \text{ days} = 12.504 \text{ years}$
- d) $\omega_E = \sqrt{\mu/a_E^3} = 1.9909 \cdot 10^{-7} \text{ rad/s}$; $\omega_{ast} = \sqrt{\mu/a_{ast}^3} = 2.1502 \cdot 10^{-7} \text{ rad/s}$
- e) Equation for stay times gives time needed to wait for favorable geometry for Hohmann return trip; since Earth and asteroid have different mean motions around Sun one of them has to orbit the Sun an integer number of orbits N more/less than the other. T : total trip time, based on Hohmann transfer "out", stay/wait time, and Hohmann transfer "back"
- f) $t_{stay} > 0 \rightarrow$ minimum value for $N = -1$ (must be an integer) $\rightarrow t_{stay} = 3.7901 \cdot 10^8 \text{ sec} = 4386.68 \text{ days} = 12.01 \text{ years} \rightarrow T = 4.0939 \cdot 10^8 \text{ sec} = 4738.33 \text{ days} = 12.97 \text{ years}$
- g) Fly faster leg to asteroid, stay for short time, fly faster leg back to Earth; will cost more energy

Question 7

- a) (6 points) Consider a single-stage rocket. Using the definitions of the payload fraction p ($p = M_{payload}/M_{begin}$) and the structural mass fraction σ ($\sigma = M_{structure}/M_{propellant}$), derive the following relation:

$$\frac{M_{begin}}{M_{end}} = \frac{1 + \sigma}{p + \sigma}$$

Here, M_{begin} is the total mass of the launcher before ignition, and M_{end} is the total mass after engine burnout.

- b) (4 points) Now consider a multi-stage launcher, where the payload fractions of the individual stages are identical (i.e. $p_{\text{tot}} = p_i^N$, where p_{tot} is the payload fraction of the entire launcher, p_i is the payload fraction of an individual stage i , and N is the total number of stages), the specific impulse is identical for the rocket engines of all stages, and the structural mass fraction σ is also the same for all stages.

Derive the following equation for the total velocity gain provided by the launcher:

$$\Delta V_{\text{tot}} = I_{sp} g_0 N \{ \ln(1 + \sigma) - \ln(\sigma + \sqrt[N]{p_{\text{tot}}}) \}$$

- c) (3 points) Compute the total velocity gain for a rocket with $I_{sp} = 400$ s, $p_{\text{tot}} = 0.02$ and $\sigma = 0.08$, for $N = 1, 2$ and 3 .
- d) (3 points) Discuss the advantages and disadvantages of replacing a single-stage launcher with a multi-stage one. What number of stages would you choose?

Data: $g_0 = 9.81$ m/s².

Answers:

- a) see sheets
b) see sheets
c) $N=1 \rightarrow \Delta V_{\text{tot}} = 9337$ m/s; $N=2 \rightarrow \Delta V_{\text{tot}} = 12436$ m/s; $N=3 \rightarrow \Delta V_{\text{tot}} = 13216$ m/s
d) Advantage: better performance (larger ΔV , more payload, combination). Disadvantage: more complexity (jettison stages, ignition) -> more risk. Also: higher development and construction costs. E.g. $N=10 \rightarrow \Delta V_{\text{tot}} = 13983$ m/s, marginal gains. In practice select $N=3$ or $N=4$, complexity not worth effort.